

Physics 240B, Spring 2008

Homework 3 Solutions

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April 7, 2008

Solutions are in general not the original work of the author.

Problem 1)

The “Ultimate efficiency” of a single-gap solar cell.

a)

The total power is the intensity times the area. Assuming all the power from the blackbody is concentrated on an area of 1 m^2 , the incident power is $I_\gamma = \sigma T^4 = 7.36 \times 10^7 \text{ W/m}^2$. This can also be done using the Planck formula to derive $\sigma = \frac{2\pi^5}{15} \frac{k_B^4}{c^2 h^3}$, arriving at the same numerical answer. So the incident power is $P_{in} = 7.36 \times 10^7 \text{ W/m}^2$.

b)

I_γ computed in (a) is much, much greater than 1000 W/m^2 . The single largest contributing factor to this is that the solid angle subtended by the earth of the sun’s outgoing power is very, very small. This solid angle is

$$d\Omega = \frac{\pi R^2}{(1A.u.)^2} = \pi \left(\frac{0.696}{149.6} \right)^2 = 6.8 \times 10^{-5} \text{ steradians} \quad (1)$$

The fraction of the intensity reaching the earth is $f_\Omega = d\Omega/\pi$, giving a total incident power of $f_\Omega I_\gamma = 1600 \text{ W/m}^2$.

c)

$$P_{out} = h\nu_g N(\nu_g) A \quad (2)$$

where $E_g = h\nu_g$ is the energy gap, $N(\nu_g)$ is the number of photons incident with energy greater than the bandgap and $A = 1 \text{ m}^2$.

For a blackbody with temperature $T (= 6000 \text{ K})$, the number of photons with energy greater than the gap is given by the Planck expression

$$N(\nu_g) = \frac{2\pi}{c} \int_{\nu_g}^{\infty} \frac{\nu^2 d\nu}{e^{\beta h\nu} - 1} \quad (3)$$

d)

The efficiency is the ratio of the output power to the incident power.

$$P_{out}/P_{in} = \frac{\frac{2\pi}{c^2} h\nu_g \int_{\nu_g}^{\infty} \frac{\nu^2 d\nu}{e^{\beta h\nu} - 1}}{\frac{2\pi}{c^2 h^3} \frac{1}{\beta^4} \frac{\pi^4}{15}} \quad (4)$$

$$\eta(x_g) = \frac{x_g}{\pi^4/15} \int_{x_g}^{\infty} \frac{x^2 dx}{e^x - 1} \quad (5)$$

where $x_g = \beta h\nu_g$. This can be solved numerically/graphically to yield $x_g = 2.17$, or $\eta = 43.8\%$. This gives a bandgap of 1.12 eV, which very nearly the gap of Si. Unfortunately Si is an indirect-gap semiconductor, so it does not form an “ultimately efficient” solar cell.

Problem 2)

The Free Energy associated with causing a superconductor to go to the normal state is:

$$\Delta F = F_N - F_S = \frac{H_c^2 V}{8\pi} \quad (\text{CGS}) \quad (6)$$

$$\Delta S = -\frac{\partial}{\partial T} \Delta F = -V \frac{H_c}{4\pi} \frac{\partial H_c}{\partial T} \quad (7)$$

Since $\frac{\partial H_c}{\partial T} < 0$, $\Delta S = S_N - S_S > 0$ as it should be.

$$\Delta C = T \frac{\partial}{\partial T} \Delta S = -\frac{VT}{4\pi} \left[\left(\frac{\partial H_c}{\partial T} \right)^2 + H_c \frac{\partial^2 H_c}{\partial T^2} \right] \quad (8)$$

By definition, $\Delta F = 0$ at T_c , therefore $H_c = 0$ at T_c , leaving

$$\Delta C(T_c) = -\frac{VT}{4\pi} \left(\frac{\partial H_c}{\partial T} \right)_{T_c}^2 \neq 0 \quad (9)$$

Note: At T_c , $\Delta S = 0 \rightarrow$ No latent heat (2nd order transition).

Problem 3)

The London Equation is:

$$\nabla^2 B - \frac{1}{\lambda^2} B = 0 \quad (10)$$

where $\lambda^2 = \frac{mc^2}{4\pi n e^2}$. From the cylindrical symmetry of the problem, $\mathbf{B} = B(\rho)\hat{\mathbf{z}}$.

Therefore,

$$\frac{\partial^2 B}{\partial \rho^2} + \frac{1}{\rho} \frac{\partial B}{\partial \rho} - \frac{1}{\lambda^2} B = 0 \quad (11)$$

This can be re-written with the dimensionless variable $x = \rho/\lambda$ as

$$\frac{\partial^2 B}{\partial x^2} + \frac{\partial B}{\partial x} - B = 0 \quad (12)$$

This is the modified Bessel Equation with index $\nu = 0$ (see Jackson, p.107-8), which has the generic solution

$$B(\rho) = aI_0\left(\frac{\rho}{\lambda}\right) + bK_0\left(\frac{\rho}{\lambda}\right) \quad (13)$$

where $I_\nu(x)$ and $K_\nu(x)$ are modified Bessel functions. Since $K_0(x) \rightarrow \infty$ as $x \rightarrow 0$, $b = 0$. At the boundary of the superconductor ($\rho = R$):

$$B(R) = aI_0\left(\frac{R}{\lambda}\right) = H_0 \quad (14)$$

$$a = \frac{H_0}{I_0\left(\frac{R}{\lambda}\right)} \quad (15)$$

So the solutions for the magnetic field are:

$$\mathbf{B} = \frac{H_0}{I_0\left(\frac{R}{\lambda}\right)} I_0\left(\frac{\rho}{\lambda}\right) \hat{\mathbf{z}}, \quad \rho \leq R \quad (16)$$

$$\mathbf{B} = H_0 \hat{\mathbf{z}}, \quad \rho \geq R \quad (17)$$

Ampere's equation gives

$$\mathbf{J} = \frac{c}{4\pi} \nabla \times \mathbf{B} = \frac{c}{4\pi} \frac{H_0}{I_0\left(\frac{R}{\lambda}\right)} \nabla \times \left(I_0\left(\frac{\rho}{\lambda}\right) \hat{\mathbf{z}} \right) \quad (18)$$

$$= \frac{c}{4\pi} \frac{H_0}{I_0\left(\frac{R}{\lambda}\right)} \frac{1}{\lambda} I_1\left(\frac{\rho}{\lambda}\right) (-\hat{\phi}) \quad (19)$$

Where $I_1(x)$ is a modified Bessel function of order $\nu = 1$.

$$\mathbf{J} = \frac{c}{4\pi} \frac{H_0}{I_0\left(\frac{R}{\lambda}\right)} \frac{1}{\lambda} I_1\left(\frac{\rho}{\lambda}\right) (-\hat{\phi}), \quad \rho \leq R \quad (20)$$

$$\mathbf{J} = 0, \quad \rho \geq R \quad (21)$$

Problem 4)

For more detail on this problem, see Kittel *Introduction to Solid State Physics* [6th ed., p.340] and Ashcroft and Mermin *Solid State Physics* [p.748].

Start with Ginzburg-Landau theory:

$$|\psi|^2 = n_s, \quad \psi = |\psi| e^{i\theta(\mathbf{r})} \quad (22)$$

and assuming pairs of charge $(-2e)$ to start:

$$\mathbf{j} = (-2e)\psi^* \mathbf{v} \psi = - \left[\frac{2e^2}{mc} \mathbf{A} + \frac{e\hbar}{m} \nabla \theta \right] |\psi|^2 \quad (23)$$

note that the velocity operator is $\mathbf{v} = \frac{1}{m} (\mathbf{p} - \frac{q}{c} \mathbf{A})$ in the presence of a field.

Now take a superconducting ring (donut/torus) and a contour C far from its surface.

We have $\oint \mathbf{j} \cdot d\mathbf{l} = 0$ which gives

$$\frac{2e^2}{mc} \oint_C \mathbf{A} \cdot d\mathbf{l} = -\frac{e\hbar}{m} \oint_C \nabla\theta \cdot d\mathbf{l} \quad (24)$$

Now

$$\oint_C \mathbf{A} \cdot d\mathbf{l} = \int_S (\nabla \times \mathbf{A}) \cdot d\mathbf{S} = \int_S \mathbf{B} \cdot d\mathbf{S} = \Phi \quad (\text{flux through } C) \quad (25)$$

and

$$-\oint_C \nabla\theta \cdot d\mathbf{l} = -\Delta\theta = 2\pi n \quad (26)$$

since ψ must be single valued. Thus:

$$|\Phi| = \frac{n\hbar c}{2e} = n\Phi_0 \quad (27)$$

where $\Phi_0 = 2.0679 \times 10^{-7}$ gauss-cm².

Kittel, ISSP has a thorough discussion of persistent currents in superconducting rings.